

## Classes of $AdS_4$ type IIA/IIB compactifications with $SU(3) \times SU(3)$ structure

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# Classes of $AdS_4$ type IIA/IIB compactifications with $SU(3) \times SU(3)$ structure

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ABSTRACT: We introduce an ansatz which allows us to solve the supersymmetry equations for warped  $\mathcal{N} = 1$   $AdS_4$  type II supergravity compactifications of general  $SU(3) \times SU(3)$  structure. As a byproduct we obtain a set of necessary conditions which every supersymmetric  $AdS_4$  vacuum should obey. The case of  $AdS_4$  compactifications of IIB on manifolds of static  $SU(2)$  structure is examined in detail. Several examples of solutions are presented. In the limit of four-dimensional Minkowski space, we present examples of supersymmetric IIB warped compactifications with partially localized NS5- and D5-branes. We also present ‘massive’ non-supersymmetric  $AdS_4 \times \mathcal{M}_6$  solutions of IIA, where  $\mathcal{M}_6$  can be any six-dimensional Einstein-Kähler manifold.

KEYWORDS: Flux compactifications, Superstring Vacua

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**1 Introduction**

$AdS_4$  vacua of type IIA string theory are examples of flux vacua in which all moduli can be stabilized at tree level, in a regime where the quantum corrections to the supergravity approximation are parametrically small. As such they appear phenomenologically promising and can serve as a starting point for the construction, upon uplifting, of metastable de Sitter vacua and models of inflation. Another strong motivation for the study of  $AdS_4$  vacua is related to the  $AdS_4/CFT_3$  duality and the recent progress in our understanding of the world-volume theory of coincident M2 branes [1, 2]. It has been observed, however, that at the moment there are many more three-dimensional superconformal field theories than there are examples of  $AdS_4$  supergravity vacua in M-theory or IIA supergravity.

All known examples to date of supersymmetric  $AdS_4$  vacua of (massive) IIA fall in the general class of rigid SU(3) solutions (an explanation of the terminology will follow shortly) given in [3]. This class includes the celebrated Nilsson-Pope  $\mathcal{N} = 6$  and  $\mathcal{N} = 1$   $AdS_4 \times \mathbb{CP}^3$

vacua [4–6] as limiting cases,<sup>1</sup> the nearly-Kähler vacua of Behrndt-Cvetič [8], as well as the vacua recently constructed by Tomasiello [9]. Finally, in [10], all previously known vacua, as well as some new ones, were constructed using left-invariant SU(3) structures on groups and cosets. On the other hand, the type IIB side has been almost entirely unexplored, perhaps due to a no-go theorem which forbids IIB  $AdS_4$  vacua with SU(3)-structure [11]. It is the purpose of this paper to go beyond the list of solutions in [10] and the analysis of [11], and take a step towards the construction of more general type II  $AdS_4$  vacua.

Supersymmetric solutions of type II supergravity of warped-product form:  $AdS_4 \times_w \mathcal{M}_6$ , where  $\mathcal{M}_6$  is the internal six-dimensional manifold, can be described in terms of two globally-defined internal spinors  $\theta_{1,2}$  specifying the spinor ansatz of the solution. These two internal spinors must be of equal norm and proportional to the warp factor, as a consequence of supersymmetry.<sup>2</sup> Hence, provided the warp factor is nowhere-vanishing, both spinors must be nowhere-vanishing. Since with each of the two internal spinors we can associate an SU(3) structure, we therefore have a *global*  $SU(3) \times SU(3)$  structure on  $\mathcal{M}_6$ . In particular it follows that there is a reduction of the structure group of  $\mathcal{M}_6$  to SU(3) or a subgroup thereof.<sup>3</sup>

The different types of solutions can be classified according to the relative angle of the two spinors. Here we follow the terminology of [14, 15], according to which we distinguish the following subcases of  $SU(3) \times SU(3)$  structure:

- *strict* SU(3) structure:  $\theta_1$  and  $\theta_2$  are parallel everywhere;
- *static* SU(2) structure:  $\theta_1$  and  $\theta_2$  are orthogonal everywhere;
- *intermediate* SU(2) structure:  $\theta_1$  and  $\theta_2$  are at a constant angle, which is neither zero nor a right angle;
- *dynamic*  $SU(3) \times SU(3)$  structure: the angle between  $\theta_1$  and  $\theta_2$  varies, possibly becoming zero or a right angle at special loci.

It was shown in [16] that there can be no IIA  $AdS_4 \times_w \mathcal{M}_6$  vacua of static SU(2) structure.<sup>4</sup> As already mentioned, there is an analogous no-go theorem in IIB forbidding  $AdS_4 \times_w \mathcal{M}_6$  vacua of strict SU(3) structure. To go beyond the static SU(2) and strict SU(3) structure cases, we must search for vacua of either dynamic  $SU(3) \times SU(3)$  or intermediate SU(2) structure.

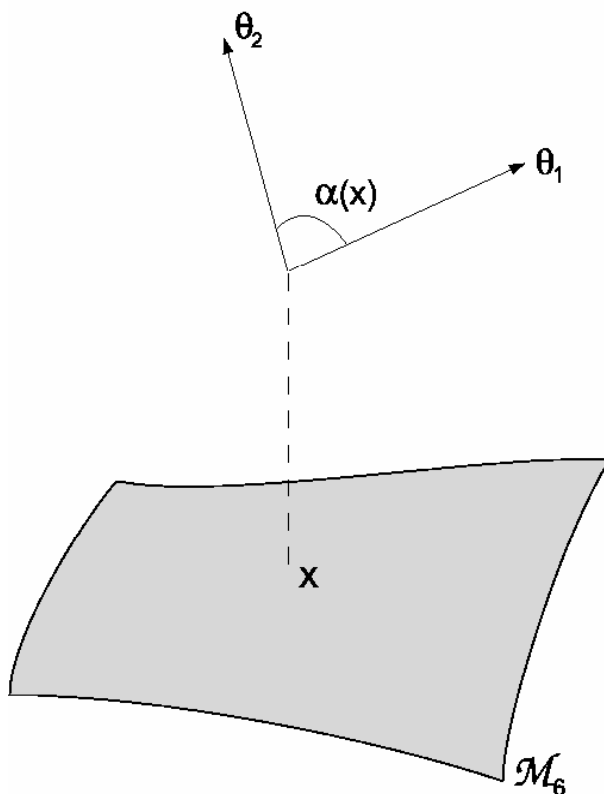
The supersymmetry equations of ten-dimensional type II supergravity for a generic global  $SU(3) \times SU(3)$ -structure ansatz can be elegantly formulated in the language of generalized geometry [17]. In searching for explicit examples of supersymmetric solutions, however, a different approach may be more promising: This is based on the observation

<sup>1</sup>The fact that the Nilsson-Pope solutions belong to the class of [3] was first pointed out in [7].

<sup>2</sup>This was first observed in [3] in the special case of rigid SU(3) structure. In the general case of  $SU(3) \times SU(3)$  structure it was first shown in the appendix of [12].

<sup>3</sup>Contrary to what is sometimes claimed in the literature, supersymmetry need not in general imply the reduction of the structure group of the internal manifold. One example is compactifications of eleven-dimensional supergravity to three-dimensional maximally-symmetric space [13].

<sup>4</sup>This no-go was subsequently generalized in [15] to include *left-invariant* intermediate SU(2) structure.



**Figure 1.** The angle  $\alpha(x)$  between the two internal spinors  $\theta_{1,2}$  is, in general, a function of the position  $x \in \mathcal{M}_6$ .

that, assuming we do not have a rigid  $SU(3)$  structure, the two  $SU(3)$  structures corresponding to each of the two internal spinors will generally intersect on a common  $SU(2)$  subgroup. In other words, we can always define a preferred *local*  $SU(2)$  structure on  $\mathcal{M}_6$ . Furthermore, we can expand all fluxes in terms of irreducible  $SU(2)$  modules, upon which the analysis of the supersymmetry conditions reduces to a set of algebraic equations for the fluxes and the torsion classes of the local structure.

The direct approach described in the preceding paragraph leads in general to cumbersome equations which cannot easily be solved, except of course in the case of rigid  $SU(3)$  structure in IIA where several solutions are known by now. In order to make progress we need to look for further simplifications. In the present paper we propose the following rather natural ansatz: we demand that the representation-theoretic content of the solution consist entirely of scalars with respect to the local  $SU(2)$  structure. In other words, in the decomposition of the various fluxes and torsion classes with respect to the local  $SU(2)$  structure, we set to zero all components which are not scalar. In the following we will refer to this as the *scalar ansatz*.

Imposing the scalar ansatz leads to considerable simplification, which enables us to explicitly solve the supersymmetry equations. The final result can be divided into two parts: (a) the part that constrains the fluxes, and (b) the part that specifies the local  $SU(2)$  struc-

ture of the internal manifold. Part (a) of the solution is given below in eqs. (2.15), (2.16) for IIA, and eqs. (2.24), (2.25) for IIB. In both cases (2.17), (2.22) hold.

There is no obstruction to solving the equations specified in part (a): they simply express some of the flux components in terms of a set of free parameters. Moreover, these equations must be satisfied by all supersymmetric solutions, not only solutions obeying the scalar ansatz. In other words, they are necessary conditions for a supersymmetric  $AdS_4$  vacuum; to our knowledge this is the first time they have been explicitly formulated.

Part (b) of the solution is given below in eq. (2.26), which is common to both IIA and IIB. Contrary to part (a) of the solution which is unobstructed, not every six-dimensional manifold will admit a local  $SU(2)$  structure obeying (2.26). Therefore, the reformulation of the supersymmetry equations in the language of the present paper provides a clear prescription for constructing new supersymmetric type II  $AdS_4$  solutions: *scan for six-dimensional manifolds which admit a local  $SU(2)$  structure obeying eq. (2.26)*.

As is well-known, supersymmetry alone is not enough to guarantee that all equations of motion are satisfied, although it goes a long way. Even in the presence of calibrated (which in the present context can be taken to mean supersymmetric) sources, there is an integrability theorem which guarantees that, provided the Bianchi identities are satisfied,<sup>5</sup> all remaining equations of motion will be automatically satisfied [18]. In general the Bianchi identities will indeed include source contributions, which may or may not admit satisfactory physical interpretation. This analysis has to be performed in addition to the analysis of the supersymmetry equations.

The remainder of the paper is organized as follows: section 2 introduces the scalar ansatz and presents the general solution to the supersymmetry equations. Section 3 contains examples of IIA solutions. In particular, section 3.1 contains examples of supersymmetric IIA solutions with smeared sources. Unfortunately these do not seem to admit a satisfactory physical interpretation. Section 3.2 contains a number of supergravity vacua of the form  $AdS_4 \times \mathcal{M}_6$ , where  $\mathcal{M}_6$  can be any six-dimensional Einstein-Kähler manifold. These solutions are shown to be non-supersymmetric, as they violate the necessary conditions of section 2.1. They are anticipated already by Romans in [19], although their existence is only mentioned very briefly in that reference (see the comment below eq. (28) of [19]).

Section 4 analyzes in detail the special case of supersymmetric  $AdS_4$  solutions of static  $SU(2)$  structure. This case is, in a sense, the analogue of the strict  $SU(3)$  case analyzed in [3], however it had not been systematically analyzed before in the literature. The complete solution to the supersymmetry equations, subject to the scalar ansatz, is given in eqs. (4.1)–(4.3) below. Section 4.1 contains two examples of solutions with smeared sources, which have appeared before in the literature, while section 4.2 contains an example with partially-localized sources, which to our knowledge is new.

The appendices A, B, C contain useful relations and many technical details of the results presented in the main text. Appendix D reviews the relation between six-dimensional Einstein-Kähler and seven-dimensional Sasaki-Einstein manifolds.

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<sup>5</sup>Here we adopt the terminology of the ‘democratic’ formalism in which the (generalized) Bianchi identities of the RR fields also include the equations of motion in the traditional sense.

**Note added:** several months after section 3.2 of the present paper was completed, we received preprint [20] which has also independently arrived at the solutions presented in that section.

## 2 Supersymmetry

In this section we introduce in detail the *scalar ansatz* referred to in the introduction, and we present the solution, under this ansatz, to the supersymmetry equations for backgrounds of the form  $AdS_4 \times_w \mathcal{M}_6$ . As a corollary we derive a set of necessary conditions (eqs. (2.15), (2.16) for IIA and eqs. (2.24), (2.25) for IIB) which must hold for any supersymmetric  $AdS_4$  vacuum — not only for vacua obeying the scalar ansatz. To our knowledge this is the first time these conditions explicitly appear in the literature.

We follow the conventions of [21], which the reader may consult for further details. We perform a four-plus-six spacetime split, according to which the ten-dimensional metric takes the warped-product form:

$$ds^2 = e^{2A(x)} ds_4^2 + g_{mn} dx^m dx^n, \quad (2.1)$$

where  $\exp A$  is the warp factor,  $ds_4^2$  is the line element of  $AdS_4$  and  $g_{mn}$  is the internal-manifold metric. The type IIA supersymmetry parameter is decomposed accordingly as:

$$\epsilon_i = \zeta \otimes \theta_i + \text{c.c.}; \quad i = 1, 2, \quad (2.2)$$

where  $\epsilon_{1,2}$  are positive-, negative-chirality ten-dimensional Majorana spinors, and  $\theta_{1,2}$  are positive-, negative-chirality six-dimensional complex spinors.  $\zeta$  is a four-dimensional positive-chirality Killing spinor obeying:

$$\nabla_\mu \zeta = \frac{1}{2} W^* \gamma_\mu \zeta^*, \quad (2.3)$$

where  $|W|$  is the inverse radius of curvature of  $AdS_4$ . Moreover we are using the democratic formalism in which the RR fluxes take the form:

$$F^{\text{tot}} = \text{vol}_4 \wedge \tilde{F} + F, \quad (2.4)$$

so that the self-duality condition reads  $\tilde{F} = \star_6 \sigma(F)$ , where  $\star_6$  is the Hodge-star on  $\mathcal{M}_6$  and  $\sigma$  reverses the order of the indices.

With these ansätze, the supersymmetry equations for type IIA/IIB can be cast in the form of a set of ‘algebraic’ equations:

$$\begin{aligned} 0 &= \not{\partial} A \theta_1 - \frac{1}{4} e^\phi \not{F} \theta_2 + e^{-A} W \theta_1^* \\ 0 &= \not{\partial} A \theta_2 - \frac{1}{4} e^\phi \gamma_7 \not{F}^\dagger \theta_1 + e^{-A} W \theta_2^* \\ 0 &= \left( \not{\partial} \phi - 2 \not{\partial} A + \frac{1}{2} \not{H} \right) \theta_1 + \frac{1}{8} e^\phi \gamma^m \not{F} \gamma_m \gamma_7 \theta_2 - 2 e^{-A} W \theta_1^* \\ 0 &= \left( \not{\partial} \phi - 2 \not{\partial} A - \frac{1}{2} \not{H} \right) \theta_2 - \frac{1}{8} e^\phi \gamma^m \not{F}^\dagger \gamma_m \theta_1 - 2 e^{-A} W \theta_2^*, \end{aligned} \quad (2.5)$$

together with a pair of ‘differential’ equations:

$$\begin{aligned} 0 &= \left( \nabla_m + \frac{1}{4} \not{H}_m \right) \theta_1 + \frac{1}{8} e^\phi \not{F} \gamma_m \gamma_7 \theta_2 \\ 0 &= \left( \nabla_m - \frac{1}{4} \not{H}_m \right) \theta_2 - \frac{1}{8} e^\phi \not{F}^\dagger \gamma_m \theta_1, \end{aligned} \tag{2.6}$$

where  $\gamma_7$  is the chirality matrix in six dimensions. Moreover  $\gamma_7 \theta_1 = \theta_1$  in both IIA/IIB, while  $\gamma_7 \theta_2 = -\theta_2$  in IIA and  $\gamma_7 \theta_2 = \theta_2$  in IIB.

**Local SU(2) structure.** For the analysis of the supersymmetry it will be useful to work with a local basis of orthogonal unimodular spinors  $\eta_{1,2}$ , with respect to which we can parameterize:

$$\theta_1 = a \eta_1; \quad \theta_2 = \begin{cases} b \eta_2^* + c^* \eta_1^* & \text{IIA} \\ b \eta_2 + c \eta_1 & \text{IIB} \end{cases}. \tag{2.7}$$

We can take  $a, b \in \mathbb{R}$ , by making use of the freedom in the definition of the phase of  $\eta_{1,2}$ , while generally  $c \in \mathbb{C}$ . This is the most general spinor ansatz, and is related to the ‘dielectric spinors’ of [14, 22, 23]. In the context of  $AdS_4$  compactifications of IIA, the two limiting cases  $b = 0$ , corresponding to rigid SU(3) structure, and  $c = 0$ , corresponding to static SU(2) structure, were considered in [3, 16] respectively. The most general spinor ansatz (2.7) has not been analyzed before in this context,<sup>6</sup> although it is of course implicit in the generalized-geometry formulation of [17].

The spinors  $\theta_{1,2}$  define a (dynamic, in general)  $SU(3) \times SU(3)$  structure, whereas the spinors  $\eta_{1,2}$  define locally a static SU(2) structure. The particular parametrization of  $\theta_{1,2}$  in terms of  $\eta_{1,2}$  above is chosen to be valid *a priori* on open patches where  $\theta_1$  is non-vanishing.<sup>7</sup> However, as already mentioned in the introduction,  $\theta_{1,2}$  are nowhere-vanishing hence this requirement is automatically satisfied (see the discussion immediately below eq. (2.22)).

Each of the two orthogonal spinors defines an SU(3) structure:

$$\begin{aligned} J_{mn}^{(r)} &:= i \eta_r \gamma_{mn} \eta_r^* \\ \Omega_{mnp}^{(r)} &:= \eta_r \gamma_{mnp} \eta_r, \end{aligned} \tag{2.8}$$

for  $r = 1, 2$ . The local static SU(2) structure  $(\tilde{J}, \omega)$  is the ‘intersection’ of these two SU(3) structures. It can be expressed in terms of  $(J^{(r)}, \Omega^{(r)}, K)$ , where  $K$  is a holomorphic one-form given by

$$K_m := \eta_2 \gamma_m \eta_1. \tag{2.9}$$

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<sup>6</sup>See [24] for certain dynamic  $SU(3) \times SU(3)$  IIA/IIB ansätze, which however do not seem to lead to solutions.

<sup>7</sup>This can be seen by ‘inverting’ (2.7) to get

$$a = |\theta_1|; \quad b = \frac{1}{|\theta_1|} \sqrt{|\theta_1|^2 |\theta_2|^2 - |\theta_1 \cdot \theta_2|^2}; \quad c^* = \frac{\theta_1 \cdot \theta_2}{|\theta_1|}.$$



As can be seen from (2.7), the additional information contained in the one-form  $(b/a)K_m$ , which e.g. in IIA is proportional to  $(\theta_2^* \gamma_m \theta_1)/|\theta_1|^2$ , can be thought of as parametrizing the deviation of the spinor ansatz from the rigid-SU(3) case. Specifically:

$$\begin{aligned} J^{(1)} &= \frac{i}{2} K \wedge K^* + \tilde{J} ; & J^{(2)} &= \frac{i}{2} K \wedge K^* - \tilde{J} \\ \Omega^{(1)} &= -i\omega \wedge K ; & \Omega^{(2)} &= i\omega^* \wedge K , \end{aligned} \quad (2.10)$$

where  $\iota_K \tilde{J}, \iota_K \omega, \iota_{K^*} \omega = 0$ . Moreover we have:

$$\omega_{mn} := i\eta_1 \gamma_{mn} \eta_2^* . \quad (2.11)$$

To analyze the content of supersymmetry, we will make repeated use of a number of additional identities satisfied by  $\eta_{1,2}$  and the various forms introduced above. These can be found in [16], whose spinor notations and conventions we follow.<sup>8</sup>

**Scalar ansatz.** The scalar ansatz proposed in the present paper consists of the following rather natural simplification: we demand that in the decomposition of the various fluxes with respect to the local SU(2) structure all components which are not scalar be set to zero.

Imposing the scalar ansatz, i.e. keeping only the scalars in the tensor decompositions given in appendix B, leads to considerable simplification upon which the various RR forms read, in form-notation:

$$\begin{aligned} e^\phi F_0 &= f_0 \\ e^\phi F_2 &= \frac{1}{8} \left( f_2 \omega^* + f_3 \tilde{J} + 2if_1 K \wedge K^* \right) + \text{c.c.} \\ e^\phi F_4 &= \frac{1}{16} g_1 \tilde{J} \wedge \tilde{J} + \frac{i}{96} \left( g_2 \omega^* + g_2^* \omega + 2g_3 \tilde{J} \right) \wedge K \wedge K^* \\ e^\phi F_6 &= f \text{vol}_6 , \end{aligned} \quad (2.12)$$

for type IIA, while:

$$\begin{aligned} e^\phi F_1 &= g_1 K + \text{c.c.} \\ e^\phi F_3 &= \frac{1}{24} \left( f_1 \omega^* + f_2 \omega + 2f_3 \tilde{J} \right) \wedge K + \text{c.c.} \\ e^\phi F_5 &= g_2 \star_6 K + \text{c.c.} , \end{aligned} \quad (2.13)$$

for type IIB. In IIA the scalars  $f, f_{0,1,3}, g_{1,3}$  are real, while  $f_2, g_2$  are complex. In IIB all five scalars  $f_{1,2,3}, g_{1,2}$  are complex. Note that in both cases the decompositions are parameterized by five complex scalar degrees of freedom. The expansion for the NSNS three-form is the same in both IIA, IIB:

$$H = \frac{1}{24} \left( h_1 \omega^* + h_2 \omega + 2h_3 \tilde{J} \right) \wedge K + \text{c.c.} , \quad (2.14)$$

where the scalars  $h_{1,2,3}$  are complex.

---

<sup>8</sup>Unlike in [16], in the present paper we do not use superspace conventions for the forms.

## 2.1 IIA solution

Plugging the expressions for the form fields (2.12), (2.14) into the algebraic supersymmetry equations (2.5) above and projecting onto the singlet of the local SU(2) structure, we obtain the following solution:

$$\begin{aligned}
 f &= -3 \operatorname{Im} \left( \frac{c}{a} W \right) e^{-A} \\
 \frac{c}{b} g_2 &= g_3 - 6i f_3 - 48i \operatorname{Im} \left( \frac{c}{a} W e^{-A} - \frac{b}{a} K \cdot \partial A \right) \\
 g_1 &= 8f_0 - \frac{2}{3} g_3 - 32 \operatorname{Re} \left( \frac{c}{a} W e^{-A} + \frac{b}{a} K \cdot \partial A \right) \\
 f_1 &= -\frac{1}{2} f_3 - \operatorname{Im} \left( \frac{c}{a} W e^{-A} - \frac{4b}{a} K \cdot \partial A \right) \\
 \frac{c}{b} f_2 &= f_3 + \frac{i}{6} g_3 + 8i \operatorname{Re} \left( \frac{c}{a} W e^{-A} + \frac{b}{a} K \cdot \partial A \right) - \frac{8ia}{b} K^* \cdot \partial A \\
 \frac{a}{b} h_3 &= \frac{3}{2} f_3 - 6 \operatorname{Im} \left( \frac{c}{a} W \right) e^{-A} - 12i \operatorname{Re} \left( \frac{c}{a} W \right) e^{-A} \\
 &\quad + 6i f_0 - \frac{i}{4} g_3 + \frac{6ia}{b} K^* \cdot \partial(3A - \phi) - \frac{12ib}{a} K^* \cdot \partial A \\
 \frac{c^*}{a} h_2 &= -\frac{i}{4} g_3 - 18i \operatorname{Re} \left( \frac{c}{a} W \right) e^{-A} - \frac{i|c|^2}{ab} \operatorname{Im} h_3 + \frac{b}{a} \operatorname{Re} h_3 \\
 &\quad + \frac{6ia}{b} \operatorname{Re} K \cdot \partial(3A - \phi) - \frac{6ib}{a} K^* \cdot \partial(3A - \phi) \\
 h_1 &= h_2^* - \frac{2ic^*}{b} \operatorname{Im} h_3 - \frac{12c^*}{b} \operatorname{Im} K \cdot \partial(3A - \phi)
 \end{aligned} \tag{2.15}$$

$$\operatorname{Im} c K \cdot \partial A = 0,$$

where we have chosen the inverse  $AdS_4$  radius  $W$ , the dilaton and warp factor  $\phi$ ,  $A$ , and the scalars  $f_0, f_3, g_3$  (see eq. (2.12)) as independent variables. Moreover we have defined  $K \cdot \partial := K^m \partial_m$ , so that  $K \cdot \partial S = \mathcal{L}_K S$ . (We use the same notation both for the one-form  $K_m dx^m$  and the vector  $K^m (\partial/\partial x^m)$  obtained by raising the covariant index with the unique metric compatible with the  $SU(3) \times SU(3)$  structure). The Romans mass is in general nonzero and enters the above equations via  $f_0 := e^\phi F_0$ .

The equations above must hold for any supersymmetric IIA  $AdS_4$  vacuum — not only for vacua obeying the scalar ansatz. To our knowledge, this is the first time they appear explicitly in the literature. In addition to these equations one would in general have a number of non-scalar equations, i.e. those which are obtained by projecting the supersymmetry equations onto irreducible representations which are not singlets under the local SU(2) structure. In the present case, these will turn out to be equivalent to (2.16), (2.26) below, as a consequence of the scalar ansatz.

In addition to the equations above, the fact that  $\eta_{1,2}$  are unimodular imposes the constraints:  $\partial(\eta_i^\dagger \eta_i) = 0$ , for  $i = 1, 2$ . There is one more constraint,  $\partial(\eta_1^\dagger \eta_2) = 0$ , which is a consequence of the orthogonality of  $\eta_{1,2}$ . Using the differential equations (2.6), it can be

seen that these three constraints are equivalent to the following:

$$\begin{aligned}
 b W e^{-A} &= \frac{c^*}{2} \operatorname{Re} K \cdot \partial \left( \log \frac{c^*}{a} + 3A - \phi \right) \\
 0 &= \operatorname{Im} K \cdot \partial \left( \log \frac{c^*}{a} + 3A - \phi \right) \\
 a &= \text{constant} \times e^{\frac{1}{2}A},
 \end{aligned}
 \tag{2.16}$$

together with:

$$dS = \frac{1}{2} K^* (K \cdot \partial S) + \frac{1}{2} K (K^* \cdot \partial S), \tag{2.17}$$

where  $S(x)$  is any one of the scalars  $A, \phi, a, b, c$ , and  $x$  is the coordinate of  $\mathcal{M}_6$ .

Before we proceed, let us make a couple of comments about eqs. (2.16), (2.17). It follows from the first two lines of (2.16) that:

$$\operatorname{Re} K W = \frac{a}{2b} e^{-2A+\phi} d \left( e^{2A-\phi} \theta_1 \cdot \theta_2 \right), \tag{2.18}$$

where we have taken footnote 7 into account together with (2.17) and the last line of (2.16). The no-go theorem of [16] then follows immediately from the above, since  $\theta_1 \perp \theta_2$  implies  $K \neq 0$  and  $W = 0$ . The more general no-go of [15] also follows similarly. Moreover, as was remarked in that reference, the way to circumvent the no-go would be to allow for  $e^{2A-\phi} \theta_1 \cdot \theta_2$  to vary over the internal manifold.

To gain insight into the meaning of equation (2.17), note that, as explained in more detail in [16],  $K$  can be used to define an almost product structure on  $\mathcal{M}_6$ . Consequently, the internal metric can locally be cast in the form:

$$ds_6^2 = \sum_{i,j=1}^4 \tilde{g}_{ij}(x) dx^i \otimes dx^j + K \otimes K^*, \tag{2.19}$$

where

$$\operatorname{Re} K = \Phi(x) \left( dx^5 + \sum_{i=1}^4 \mathcal{A}_i(x) dx^i \right); \quad \operatorname{Im} K = \Psi(x) \left( dx^6 + \sum_{i=1}^4 \mathcal{B}_i(x) dx^i \right). \tag{2.20}$$

Since  $\tilde{g}_{ij}, \Phi, \Psi, \mathcal{A}_i, \mathcal{B}_i$  depend in general on all coordinates of  $\mathcal{M}_6$ , it follows that (2.19) is not in general a fibration. Condition (2.17) can then locally be rewritten as:

$$\frac{\partial}{\partial x^i} S = 0; \quad i = 1, \dots, 4. \tag{2.21}$$

Finally, in order to allow for  $\text{AdS}_4$  solutions,  $W \neq 0$ , it turns out that  $a, b, c$  must satisfy the following relation:

$$a^2 = b^2 + |c|^2. \tag{2.22}$$

Equivalently, the measures of the two spinors  $\theta_{1,2}$  must be equal:

$$|\theta_1|^2 = |\theta_2|^2. \tag{2.23}$$

As already mentioned in the introduction, it follows from (2.22), or equivalently (2.23), and the last equation in (2.16) that  $\theta_{1,2}$  must be nowhere-vanishing. We therefore have a *globally* well-defined  $SU(3) \times SU(3)$  structure on  $\mathcal{M}_6$ .

It is straightforward to verify that the results of [16] are recovered in the  $c \rightarrow 0$  limit, which corresponds to the static  $SU(2)$  case. The limit  $b \rightarrow 0$ , which corresponds to the strict  $SU(3)$  case [3], can also be taken but is slightly more subtle, as in this limit the irreducible representations which appear in the tensor decompositions of the various fields, have to be taken with respect to the  $SU(3)$  structure.

## 2.2 IIB solution

Proceeding similarly to the IIA case, taking eqs. (2.13), (2.14) into account, the algebraic supersymmetry equations (2.5) can be solved to give:

$$\begin{aligned}
 f_1 &= 12i \left\{ \frac{c}{b}(g_1 + ig_2) + \frac{c}{a}W^*e^{-A} + \left(\frac{2a}{b} - \frac{b}{a}\right)K^* \cdot \partial A \right\} \\
 f_2 &= 12i \left\{ -\frac{c}{a}W^*e^{-A} + \frac{b}{a}K^* \cdot \partial A \right\} \\
 f_3 &= 12i \left\{ \frac{1}{2}(g_1 + ig_2) + \frac{b}{a}W^*e^{-A} - \frac{c}{a}K^* \cdot \partial A \right\} \\
 h_1 &= 12i \left\{ \left(\frac{a}{b} - \frac{b}{2a}\right)(g_1 - ig_2) - W^*e^{-A} + \frac{c}{b}K^* \cdot \partial(2A - \phi) \right\} \\
 h_2 &= 12i \left\{ \frac{b}{2a}(g_1 - ig_2) - W^*e^{-A} \right\} \\
 h_3 &= 6i \left\{ -\frac{c}{a}(g_1 - ig_2) + K^* \cdot \partial(2A - \phi) \right\} \\
 \text{Re } c &= 0 .
 \end{aligned} \tag{2.24}$$

Note that the solution leaves the complex scalars  $g_1, g_2$  unconstrained. Moreover, the constraints  $\partial(\eta_i^\dagger \eta_j) = 0$  imply:

$$\begin{aligned}
 b W e^{-A} &= \frac{c}{3} K \cdot \partial \left( \phi - 4A - \log \frac{|c|}{a} \right) + \frac{2ia}{3} g_2^* \\
 a &= \text{constant} \times e^{\frac{1}{2}A} .
 \end{aligned} \tag{2.25}$$

The equations above must hold for any supersymmetric IIB  $AdS_4$  vacuum — not only for vacua obeying the scalar ansatz. To our knowledge, this is the first time they appear explicitly in the literature. In addition, eqs. (2.17), (2.22) hold in the present case as well.

## 2.3 Local $SU(2)$ structure

The local  $SU(2)$  structure of the internal manifold, encoded in the action of the exterior differential on  $(K, \tilde{J}, \omega)$ , can be read off using the differential supersymmetry equations (2.6) as explained in appendix C. More specifically, for both IIA and IIB we can give the following

compact expressions:

$$\begin{aligned}
 dK &= K^* \wedge K \left\{ \frac{1}{2}(K^*)_1 - \frac{1}{2}(K)_1^* - (K^*K)_2 + (KK)_2^* \right\} \\
 &\quad + \omega \{-4(\omega)_2\} + \omega^* \{2(\tilde{J})_1\} + \tilde{J} \{-2(\tilde{J})_2 - 4(\omega)_1\} \\
 d\tilde{J} &= K \wedge \omega \left\{ -2(K^*K)_1^* - i(\tilde{J})_1^* \right\} + K \wedge \omega^* \{-2(KK)_1 - 2i(\omega)_2^*\} \\
 &\quad + K \wedge \tilde{J} \{2i(\omega)_1^* - i(\tilde{J})_2^*\} + \text{c.c.} \\
 d\omega &= K \wedge \tilde{J} \{4(KK)_1 + 4i(\omega)_2^*\} + K^* \wedge \tilde{J} \{4(K^*K)_1 - 2i(\tilde{J})_1\} \\
 &\quad + K \wedge \omega \left\{ \frac{1}{2}(K)_1 - \frac{1}{2}(K^*)_1^* - (K^*K)_2^* + (KK)_2 - 2i(\tilde{J})_2^* \right\} \\
 &\quad + K^* \wedge \omega \left\{ \frac{1}{2}(K^*)_1 - \frac{1}{2}(K)_1^* + (K^*K)_2 - (KK)_2^* - 4i(\omega)_1 \right\},
 \end{aligned} \tag{2.26}$$

which can be derived from (C.7), (C.8), (C.9) with the use of (C.17). All coefficients on the right-hand sides above are known and are explicitly given in eqs. (C.3)–(C.6). Since the geometry is determined by the local SU(2) structure, eq. (2.26) fixes the geometry in terms of the flux parameters.

Note that the local SU(2) structure can also be specified either by the triplet  $(K, J^{(1)}, \Omega^{(1)})$ , or, equivalently,  $(K, J^{(2)}, \Omega^{(2)})$ . In the former case (2.26) would have to be replaced by the expression for  $dK$  (the first of the equations above) together with the expression for the torsion classes, given in (C.19), for the SU(3) structure corresponding to  $(J^{(1)}, \Omega^{(1)})$ . As already remarked below (2.9), the additional information contained in the one-form  $(b/a)K$  can be thought of as parametrizing the deviation of the spinor ansatz from the rigid-SU(3) case.

**In summary:** *for a supersymmetric background of the form  $AdS_4 \times \mathcal{M}_6$ , the internal manifold  $\mathcal{M}_6$  is specified by a local SU(2) structure  $(K, \tilde{J}, \omega)$  obeying (2.26); the fluxes are given by (2.15), (2.16) in IIA, and by (2.24), (2.25) in IIB; in both cases (2.17), (2.22) hold.*

### 3 IIA examples

The reformulation of the supersymmetry equations in the present language readily suggests a strategy for a systematic search for solutions: Given an SU(3)-structure manifold  $\mathcal{M}_6$  choose a family of triplets  $(K^\lambda, \tilde{J}^\lambda, \omega^\lambda)$  on it, where  $\lambda$  parameterizes the family; impose eqs. (2.26) in order to restrict  $\lambda$ ; if a solution exists on  $\mathcal{M}_6$ , read off the fluxes using (2.15), (2.16). The following examples will illustrate this method for type IIA. In the next section we will consider the case of static SU(2) structure in IIB.

#### 3.1 Examples with smeared sources

The following is a simple solution of the supersymmetry equations. Let us demand that  $d\omega$  should not contain any  $K \wedge \tilde{J}$ ,  $K^* \wedge \tilde{J}$  terms. This can be seen from (2.26) to automatically

imply that  $d\tilde{J}$  contains only  $K \wedge \tilde{J}$ ,  $K^* \wedge \tilde{J}$  terms. In addition, we demand that  $dK$  be proportional to  $K^* \wedge K$ . We will also assume that  $\text{Im}c \neq 0$ . As we can see from (2.15), this implies  $K \cdot \partial A = 0$ . Taking the constraints (2.16) into account, the aforementioned conditions imply:

$$\begin{aligned}
 f_0 &= \frac{4b^2 + 5|c|^2}{2ab}C; & f_1 &= 0; & f_2 &= \frac{2ic^*}{a}C; & f_3 &= 0 \\
 g_1 &= \frac{12a^2 + 4b^2}{ab}C; & g_2 &= \frac{36c^*}{a}C; & g_3 &= \frac{36|c|^2}{ab}C \\
 h_1 &= 0; & h_2 &= -\frac{12ic}{b}C; & h_3 &= 6iC \\
 \text{Im}(cW) &= 0; & \text{Im}K \cdot \partial\phi &= 0; & \text{Re}K \cdot \partial\phi &= C; & a, b, c, A &= \text{constant},
 \end{aligned} \tag{3.1}$$

where we have introduced the real constant  $C := -(2b/c^*)e^{-A}W$ . It readily follows from the above that we have an intermediate  $\text{SU}(2)$  structure.

In form notation the fluxes read:

$$\begin{aligned}
 H &= \frac{i}{2}C \left( \tilde{J} - \frac{c}{b}\omega^* \right) \wedge K + \text{c.c.} \\
 e^\phi F_0 &= \frac{4b^2 + 5|c|^2}{2ab}C; & e^\phi F_2 &= -\frac{ic}{4a}C^*\omega + \text{c.c.} \\
 e^\phi F_4 &= \frac{3a^2 + b^2}{4ab}C\tilde{J} \wedge \tilde{J} + \frac{3i}{4a} \left( \frac{|c|^2}{b}C\tilde{J} + \text{Re}(cC^*\omega) \right) K \wedge K^*.
 \end{aligned} \tag{3.2}$$

Furthermore, we can compute the local structure from (2.26):

$$\begin{aligned}
 d\text{Re}K &= 0; & d(e^\phi \text{Im}K) &= 0 \\
 d(e^{-\phi}\omega) &= 0; & d(e^{-\phi}\tilde{J}) &= 0.
 \end{aligned} \tag{3.3}$$

The above relations imply that  $K$  can be written as  $K = d\varphi + ie^{-\phi}d\chi$  for some local coordinates  $\varphi, \chi$ . It then follows from (3.1) that the dilaton is given by

$$\phi = C(\varphi - \varphi_0), \tag{3.4}$$

for some constant  $\varphi_0$ . Moreover, as can be seen from (C.10), the two-forms  $(e^{-\phi}\omega), (e^{-\phi}\tilde{J})$  define a four-dimensional Calabi-Yau manifold, i.e. a  $K3$  surface. The metric of the six-dimensional internal manifold can therefore be written as:

$$ds_6^2 = e^\phi ds_{K3}^2 + d\varphi^2 + e^{-2\phi}d\chi^2, \tag{3.5}$$

where  $ds_{K3}^2$  is the metric of the  $K3$  surface. Note that  $\varphi, \chi$  parameterize a two-dimensional hyperbolic space  $H_2$ .

Although the supersymmetry equations can be solved in the way described above, it is not difficult to see that the sourceless Bianchi identities cannot be satisfied for all form fields. In particular, negative-tension (non-localized) sources must be added, which is physically unsatisfactory. Although we will not list the details here, similar solutions of the

supersymmetry equations (but not of the sourceless Bianchi identities) can be achieved by taking the internal manifold to be a nilmanifold. It is possible that performing a systematic scan of the nilmanifolds, something which we have not done, would yield supersymmetric solutions which also satisfy the sourceless Bianchi identities.

**Constant warp factor, dilaton.** In the case of constant dilaton and warp factor, a simple way to solve (2.16), (2.22) is by making the following ansatz:

$$b = a \cos \varphi; \quad c = ae^{i\delta} \sin \varphi; \quad W = |W|e^{-i\delta}; \quad \phi, A, a, \delta = \text{constant}, \quad (3.6)$$

where we have parameterized:

$$\text{Re}K = \frac{e^A}{2|W|}d\varphi + \mathcal{A}, \quad (3.7)$$

for some co-ordinate  $\varphi$  and a one-form  $\mathcal{A}$  such that  $\iota_{\partial/\partial\varphi}\mathcal{A} = 0$ . In order to see that (3.6) is indeed a solution of (2.16), (2.22), note that (3.7) implies  $\text{Re}K \cdot \partial = 2|W|e^{-A}\partial/\partial\varphi$ .

### 3.2 Examples without sources

We will now consider a certain class of IIA compactifications of the form  $AdS_4 \times \mathcal{M}_6$ , where  $\mathcal{M}_6$  can be any Einstein-Kähler manifold. We will allow for non-zero Romans mass, therefore these compactifications do not, in general, admit an eleven-dimensional lift. These solutions were anticipated by Romans in [19] (see also [25]), although their existence was only mentioned very briefly in that reference (cf. the comment below eq. (28) of [19]).

For non-vanishing Romans mass these solutions will be shown, at the end of the present section, to be non-supersymmetric, as they do not obey the necessary supersymmetry conditions of section 2.1. On the other hand, for vanishing Romans mass we have an enhancement of supersymmetry, and the solutions fall within the class of the supersymmetric solutions of [3]. Using the known results, summarized in section D, relating six-dimensional Einstein-Kähler manifolds to seven-dimensional Sasaki-Einstein manifolds, for vanishing Romans mass these solutions lift to the well-known supersymmetric M-theory solutions of Freund-Rubin type of the form  $AdS_4 \times \mathcal{M}_7$ , where  $\mathcal{M}_7$  is Sasaki-Einstein.

We take the ten-dimensional metric to be of the form:

$$ds^2 = ds^2(AdS_4) + ds^2(\mathcal{M}_6), \quad (3.8)$$

i.e. a direct (not warped) product  $AdS_4 \times \mathcal{M}_6$ . Moreover, we take the NSNS three-form to vanish,  $H = 0$ , and the dilaton to be constant. The RR fields are given by:

$$F_0 = \alpha; \quad F_2 = \beta J; \quad F_4 = \frac{1}{2}\gamma J^2; \quad F_6 = \frac{1}{6}\delta J^3, \quad (3.9)$$

where  $J$  is the Kähler form on  $\mathcal{M}_6$ , and  $\alpha, \dots, \delta \in \mathbb{R}$ . After imposing the self-duality condition, see below eq. (2.4), the RR fluxes can be written more conventionally as:

$$F_0^{\text{tot}} = \alpha; \quad F_2^{\text{tot}} = \beta J; \quad F_4^{\text{tot}} = \frac{1}{2}\gamma J^2 + \delta \text{vol}_4. \quad (3.10)$$

The following calculations are very similar to section 11.4 of [21], so here we will simply state the results.

The NSNS Bianchi identity,  $dH = 0$ , is trivially satisfied for this ansatz. Similarly, the generalized Bianchi identities for the RR fields,  $d_H F = 0$ , (which in the conventional type II supergravity formulation correspond to both the Bianchi identities and the equations of motion) are also automatically satisfied by virtue of the closure of the Kähler form,  $dJ = 0$ . It remains to examine the NS-sector equations of motion. The  $H$ -field equation of motion reduces to

$$\alpha\beta + 2\beta\gamma + \gamma\delta = 0 . \quad (3.11)$$

The dilaton equation reads:

$$|W|^2 - \frac{5}{8}\omega^2 = 0 , \quad (3.12)$$

where  $W \in \mathbb{C}$ ,  $\omega \in \mathbb{R}$  are related to the curvature of  $AdS_4$ ,  $\mathcal{M}_6$  via

$$R_{\mu\nu} = -3g_{\mu\nu}|W|^2, \quad R_{mn} = \frac{5}{4}\omega^2 g_{mn} , \quad (3.13)$$

respectively. Finally, the external and internal Einstein equations read:

$$|W|^2 - \frac{1}{12}(\alpha^2 + 3\beta^2 + 3\gamma^2 + \delta^2) = 0 \quad (3.14)$$

and

$$5\omega^2 + \alpha^2 + \beta^2 - \gamma^2 - \delta^2 = 0 , \quad (3.15)$$

respectively.

The full set of supergravity equations of motion above can be seen to admit three infinite classes of solutions. In each of these three classes, the constants  $|W|$ ,  $\omega$  can be solved for in terms of the real parameters  $\alpha, \dots, \delta$  using (3.12), (3.14). Moreover we have:

*First solution:*

$$\beta = \gamma = 0 ; \quad \delta = \pm\sqrt{5}\alpha . \quad (3.16)$$

*Second solution:*

$$\alpha = \pm\frac{7}{5\sqrt{5}}\beta ; \quad \gamma = \pm\frac{1}{\sqrt{5}}\beta ; \quad \delta = -\frac{17}{5}\beta . \quad (3.17)$$

*Third solution:*  $\beta^2 \geq 3\gamma^2$  and

$$\alpha = \gamma \frac{-2\beta^2 \pm \sqrt{(\beta^2 - 3\gamma^2)(9\beta^2 + 5\gamma^2)}}{\beta^2 - 5\gamma^2} ; \quad \delta^2 = 5\alpha^2 + 9\beta^2 + 3\gamma^2 . \quad (3.18)$$

The Nilsson-Pope ‘Hopf-fibration’ solution [4] is a subset of the third solution above, and is obtained upon setting the Romans mass to zero,  $\alpha = 0$ . In this case we obtain:

*Hopf-fibration solution:*

$$\alpha = \gamma = 0 ; \quad \delta = \pm 3\beta . \quad (3.19)$$

Comparing with (3.10) we see that this solution corresponds to a Freund-Rubin ansatz,  $F_4^{\text{tot}} \propto \text{vol}_4$ , with  $F_0^{\text{tot}} = 0$  and  $F_2^{\text{tot}} \propto J$ .



**Supersymmetry.** Let us now consider the supersymmetry of the solutions above. Imposing  $H = 0$  in addition to eqs. (2.15) implies:

$$e^\phi F_0 = -\text{Re}\left(\frac{c}{a}W\right)e^{-A}; \quad g_2 = -\frac{72b}{a}We^{-A}. \quad (3.20)$$

Since  $\phi, A, F_0$  are constant, it follows from the above that  $c/a$  is constant. If  $b \neq 0$ , the first equation in (2.16) then implies that  $W = 0$  and consequently  $AdS_4$  decompactifies to flat Minkowski space. If on the other hand  $b = 0$ , the situation reduces to the rigid  $SU(3)$  case, as follows from eq. (3.5). The solution then falls within the class of supersymmetric  $AdS_4$  solutions of [3], from which it follows that supersymmetry enforces  $F_0 = 0$ .

**In summary:** for nonzero Romans mass, the solutions presented in this section are not supersymmetric, as they violate the necessary conditions of section 2.1. For vanishing Romans mass there is an enhancement of supersymmetry, and these solutions fall within the class of the supersymmetric solutions of [3].

#### 4 Static $SU(2)$ structure in IIB

It has been known for some time that static  $SU(2)$ -structure compactifications to  $AdS_4$  are not allowed in IIA [16]. There is a IIB counterpart of this no-go, forbidding strict  $SU(3)$ -structure compactifications to  $AdS_4$  in IIB [11]. However, static  $SU(2)$ -structure compactifications to  $AdS_4$  are allowed in IIB. In this case we have  $a = \pm b, c = 0$ , cf. eq. (2.7), and eqs. (2.24), (2.25) simplify considerably to:

$$\begin{aligned} f_1 = f_2 &= \pm 12iK^* \cdot \partial A \\ f_3 &= 6i(g_1 + ig_2) \pm 12iW^*e^{-A} \\ h_1 = h_2 &= \pm 6i(g_1 - ig_2) - 12iW^*e^{-A} \\ h_3 &= 6iK^* \cdot \partial(2A - \phi) \end{aligned} \quad (4.1)$$

and

$$g_2 = \pm \frac{3i}{2}W^*e^{-A}; \quad a = \pm b = \text{constant} \times e^{\frac{1}{2}A}, \quad (4.2)$$

respectively. The  $SU(2)$  structure, which can be read off off (2.26), (C.5), (C.6), can be put in the form:

$$\begin{aligned} dK_0 &= -2W\text{Im}\omega_0 \\ d\tilde{J}_0 &= \mp g_0 K_0 \wedge \text{Re}\omega_0 + \text{c.c.} \\ d\omega_0 &= \pm g_0 K_0 \wedge \tilde{J}_0 + \text{c.c.}, \end{aligned} \quad (4.3)$$

where we have set:

$$K_0 := e^{3A-\phi}K; \quad \tilde{J}_0 := e^{2A-\phi}\tilde{J}; \quad \omega_0 := e^{2A-\phi}\omega; \quad g_0 := \frac{1}{2}e^{\phi-3A} \left( g_1 \mp \frac{5}{2}W^*e^{-A} \right). \quad (4.4)$$

Consistency requires that  $d^2$  should annihilate  $K_0, \tilde{J}_0, \omega_0$ , which is guaranteed provided  $g_0$  is ‘holomorphic’ (cf. the discussion around eq. (2.19)):

$$dg_0 = \frac{1}{2}K(K^* \cdot \partial g_0) . \tag{4.5}$$

A special solution of the above is  $g_0 = \text{constant}$ .

**Constant warp factor, dilaton.** A further simplification to eqs. (4.1)–(4.3) would be to assume constant dilaton and warp factor. Setting  $\phi = A = 0$  and demanding that  $g_1$  be holomorphic, i.e. that it should satisfy the analogue of (4.5), it is now straightforward to examine the Bianchi identities and equations of motion for all the form fields. Imposing  $dH = 0$  (i.e. demanding the absence of NS5 brane sources) implies:

$$\text{Re}(g_1 W) = \frac{1}{3} \left( |g_1|^2 + \frac{5}{4}|W|^2 \right) , \tag{4.6}$$

as follows from (2.14), (4.3). Moreover we find a source (D7 branes/O7 planes) for the Bianchi identity of  $F_1$ :

$$dF_1 = -\frac{4}{3} \left( |g_1|^2 + \frac{5}{4}|W|^2 \right) \text{Im} \omega . \tag{4.7}$$

The source above corresponds to net orientifold charge. Note that demanding the absence of D7/O7 sets the cosmological constant to zero.

In addition there is a potential source (D5 branes/O5 planes), which vanishes for special values of  $g_1$ , for the Bianchi identity of  $F_3$ :

$$dF_3 + H \wedge F_1 = i \left( |g_1|^2 - \frac{5}{4}|W|^2 \right) \text{Re} \omega \wedge K \wedge K^* . \tag{4.8}$$

There is a net orientifold charge for  $|g_1| \geq \sqrt{5}/2|W|$ . All other Bianchi’s and equations of motion for the form-fields are automatically satisfied. It is then guaranteed by the integrability theorem of [18], which generalizes the theorems of [3, 26] to include calibrated sources, that all remaining equations of motion are automatically satisfied.

#### 4.1 Examples with smeared sources

In the following we will discuss two examples of supersymmetric IIB  $AdS_4$  compactifications solving eqs. (4.1)–(4.3). Both of these examples, which have been mentioned before in the literature, contain sources smeared in the internal space.

**Nilmanifold 5.1.** This example, where we take the internal six-dimensional manifold to be the nilmanifold 5.1, was first mentioned in [27] and further examined in [15]. The nilmanifold 5.1 can be defined by specifying a coframe  $e^i, i = 1, \dots, 6$ , such that:

$$de^i = 0, \quad i = 1, \dots, 5 ; \quad de^6 = e^{12} + e^{34} , \tag{4.9}$$

where  $e^{ij} := e^i \wedge e^j$ . Let us set  $A, \phi = 0$  for simplicity. Moreover, assuming  $a = +b$ , let us take

$$g_1 = \frac{5}{2}W^* , \tag{4.10}$$

so that  $g_0 = 0$ , by virtue of (4.4). Eqs. (4.3) are then satisfied, provided we identify:

$$\begin{aligned} K &= -2W e^6 + i e^5 \\ \tilde{J} &= e^{13} - e^{24} \\ \omega &= (i e^1 + e^3) \wedge (i e^4 + e^2) . \end{aligned} \tag{4.11}$$

This solution contains (smeared) O5/O7 sources, as can be seen by computing the right-hand-sides of eqs. (4.7), (4.8) above taking (4.10) into account.

$\mathbf{T}^{1,1} \times \mathbf{S}^1$ . In this example, which was first mentioned in [10], we take the internal six-dimensional manifold to be the product  $T^{1,1} \times S^1$ . The total six-dimensional manifold admits a coset structure, described in section 4.6 of ref. [10], to which the reader is referred for further details.<sup>9</sup> As in the previous case, we can describe the internal manifold by specifying a coframe  $e^i$ ,  $i = 1, \dots, 6$ . The action of the exterior differential on the coframe is determined by the structure constants of the coset. As before, let us set  $A$ ,  $\phi = 0$ . Eqs. (4.3) are then satisfied, provided we identify:

$$\begin{aligned} K &= 2W e^3 + i e^6 \\ \tilde{J} &= -e^{14} + e^{25} \\ \omega &= -i(i e^1 + e^4) \wedge (i e^2 - e^5) . \end{aligned} \tag{4.12}$$

In addition we must take  $g_0 = -1/2W \in \mathbb{R}$ , so that:

$$g_1 = -\frac{1}{W} + \frac{5}{2}W . \tag{4.13}$$

As in the previous example, this solution contains (smeared) O5/O7 sources. It also generally contains (smeared) NS5-brane sources, which vanish for the special value:  $W = \pm 1/\sqrt{2}$ , as can be seen from (4.6), (4.13).

## 4.2 Examples with partially localized sources

Taking the limit to four-dimensional Minkowski space ( $W \rightarrow 0$ ), we will now discuss a class of supersymmetric IIB warped compactifications solving eqs. (4.1), (4.3). These examples contain spacetime-filling NS5 and/or D5 branes partially localized in the internal space.

Let us take  $g_1 = g_2 = 0$ , so that  $W = 0$ , in which case the external space becomes  $\mathbb{R}^{1,3}$ . It follows from (4.3) that the two-forms  $(e^{2A-\phi}\omega)$ ,  $(e^{2A-\phi}\tilde{J})$  are closed, and therefore define a four-dimensional Calabi-Yau manifold, i.e. a  $K3$  surface. It also follows from (4.3) that the one-form  $(e^{3A-\phi}K)$  is closed. We can therefore take it to be equal to  $\mathcal{D}z := dz + \mathcal{A}$ , where  $\mathcal{A}$  is a flat connection on  $K3$ , and  $z$  is a complex coordinate of a  $T^2$ .

The metric of the six-dimensional internal manifold can therefore be written as:

$$ds_6^2 = e^{2\phi-6A} |\mathcal{D}z|^2 + e^{\phi-2A} ds_{K3}^2 , \tag{4.14}$$

---

<sup>9</sup>The present case corresponds to the  $b = 0$  embedding described in eqs. (4.36,4.37) of ref. [10].

where  $ds_{K3}^2$  is the metric of the  $K3$  surface. Moreover, the non-zero fluxes can be read off off (2.13), (2.14):

$$\begin{aligned}
 F_3 &= -\frac{i}{2} \frac{\partial}{\partial z} (e^{-2A}) \operatorname{Re} \omega_0 \wedge \mathcal{D}z + \text{c.c.} \\
 H &= -\frac{i}{2} \frac{\partial}{\partial z} (e^{\phi-2A}) \tilde{J}_0 \wedge \mathcal{D}z + \text{c.c.} .
 \end{aligned}
 \tag{4.15}$$

As is now straightforward to compute, there will, in general, be source-terms in the Bianchi identities for the above form-fields, signalling the presence of NS5 and/or D5 branes. Indeed we find:

$$\begin{aligned}
 dF_3 &= \frac{i}{2} \frac{\partial^2}{\partial z \partial z^*} (e^{-2A}) \operatorname{Re} \omega_0 \wedge \mathcal{D}z \wedge (\mathcal{D}z)^* + \text{c.c.} \\
 dH &= \frac{i}{2} \frac{\partial^2}{\partial z \partial z^*} (e^{\phi-2A}) \tilde{J}_0 \wedge \mathcal{D}z \wedge (\mathcal{D}z)^* + \text{c.c.} .
 \end{aligned}
 \tag{4.16}$$

Taking the functions  $e^{-2A}$ ,  $e^{\phi-2A}$  to be harmonic on  $T^2$  ensures that the source-terms on the right-hand sides above are localized on  $T^2$ .

To complete the discussion of these solutions, one can also show that all remaining Bianchi identities and equations of motion for the form fields are satisfied for the system of fluxes given in (4.15). As already remarked, the integrability theorem of [18] then guarantees that all remaining equations of motion will be automatically satisfied.

## 5 Conclusions

The scalar ansatz introduced in the present paper allowed us to explicitly solve the supersymmetry equations of type II supergravity. The ‘algebraic part’ of the solution is given by eqs. (2.15), (2.16) for IIA, and eqs. (2.24), (2.25) for IIB. Moreover, these are necessary conditions which every supersymmetric  $AdS_4$  solution should obey — not only the solutions satisfying the scalar ansatz. In addition, eqs. (2.17), (2.22) must be imposed in both cases.

As already pointed out in the introduction, the algebraic part of the solution is unobstructed, as it simply expresses certain flux components in terms of a set of free parameters. The ‘differential part’ of the solution is given in eq. (2.26), and specifies the local  $SU(2)$  structure of the internal manifold. The main message of the present paper is therefore that: *in order to construct new supersymmetric  $AdS_4$  compactifications of type II supergravity, it suffices to find six-dimensional manifolds which admit a local  $SU(2)$  structure obeying eq. (2.26).* A natural direction for further study would be to systematically scan different classes of manifolds for that purpose.

Solutions of the supersymmetry equations will in general contain sources. The source content of a solution is revealed by studying the Bianchi identities of the form fields. As we have seen in the examples presented here, the sources present in a solution may or may not admit a satisfactory physical interpretation. At least one need not worry about the remaining equations of motion: thanks to the integrability theorem of [18], we know that these will be automatically satisfied.

The case of  $AdS_4$  compactifications of IIB on manifolds of static  $SU(2)$  structure is, in some sense, the analogue of the well-known strict- $SU(3)$  case in IIA. Nevertheless, it had not been systematically studied before. In section 4 we examined this case in detail. In particular, eqs. (4.1), (4.2) are necessary conditions that every supersymmetric  $AdS_4$  solution of static  $SU(2)$  structure should obey. The examples of solutions presented in section 4.1, had already appeared in the literature in [10, 27], whereas to our knowledge the example of section 4.2 is new. The latter is obtained in the limit of four-dimensional Minkowski space, and contains partially localized NS5- and D5-branes. It is perhaps worth noting that this example does not fall into the GKP class [29].

The nonsupersymmetric solutions presented in section 3.2 were anticipated by Romans already in [19], although they only received a brief mention in that reference. As we have seen, these solutions naturally fall into three distinct classes, eqs. (3.16)–(3.18), the last of which can be thought of as a deformation of the Nilsson-Pope solution. The  $CFT_3$  dual of the latter class was recently considered in [20]. It would be interesting to examine whether a  $CFT_3$  dual can also be constructed for the other two classes.

### A Useful relations

In this section we list the following relations which are useful in deriving the supersymmetry conditions of section 2.1. For a more complete list the reader may consult [16].

$$\begin{aligned}
 H\eta_1^* &= -\frac{i}{3}h_2\eta_1 - \frac{i}{6}h_3K_m\gamma^m\eta_1^* \\
 H\eta_2^* &= \frac{i}{3}h_3\eta_1 - \frac{i}{6}h_1K_m\gamma^m\eta_1^* \\
 H_m\eta_1 &= \frac{i}{6}(h_3K_m + h_3^*K_m^*)\eta_1 + \frac{1}{12}(2h_2^*\tilde{J}_{mn} - h_3^*\omega_{mn} - ih_1K_mK_n - ih_2^*K_m^*K_n)\gamma^n\eta_1^* \\
 H_m\eta_2 &= \frac{i}{6}(h_2K_m + h_1^*K_m^*)\eta_1 + \frac{1}{12}(-2h_3^*\tilde{J}_{mn} - h_1^*\omega_{mn} + ih_3K_mK_n + ih_3^*K_m^*K_n)\gamma^n\eta_1^*
 \end{aligned} \tag{A.1}$$

and

$$\begin{aligned}
 e^\phi F\eta_1 &= \left\{ f_0 - \frac{1}{8}g_1 - \frac{1}{12}g_3 + i\left(f + f_1 + \frac{1}{2}f_3\right) \right\} \eta_1 + \left(-\frac{i}{4}f_2 + \frac{1}{24}g_2\right) K_m\gamma^m\eta_1^* \\
 e^\phi F\eta_2 &= \left(\frac{i}{2}f_2^* - \frac{1}{12}g_2^*\right) \eta_1 - \frac{1}{2} \left\{ f_0 - \frac{1}{8}g_1 + \frac{1}{12}g_3 + i\left(f + f_1 - \frac{1}{2}f_3\right) \right\} K_m\gamma^m\eta_1^* \\
 e^\phi F\gamma_m\eta_1^* &= \left(-\frac{i}{2}f_2^* + \frac{1}{12}g_2^*\right) K_m^*\eta_1 \\
 &+ \left\{ \left(f_1 - f + \frac{i}{8}g_1 + if_0\right) \tilde{J}_{mn} + \frac{1}{2} \left( if + if_1 - \frac{i}{2}f_3 - \frac{1}{8}g_1 + \frac{1}{12}g_3 + f_0 \right) K_m^*K_n \right\} \gamma^n\eta_1^* \\
 e^\phi F\gamma_m\eta_2^* &= \left( if + if_1 + \frac{i}{2}f_3 - \frac{1}{8}g_1 - \frac{1}{12}g_3 + f_0 \right) K_m^*\eta_1 \\
 &+ \frac{1}{2} \left\{ \left(f_1 - f + \frac{i}{8}g_1 + if_0\right) \omega_{mn} + \left(-\frac{i}{2}f_2 + \frac{1}{12}g_2\right) K_m^*K_n \right\} \gamma^n\eta_1^*,
 \end{aligned} \tag{A.3}$$

$$\begin{aligned}
 e^\phi F\gamma_m\eta_1^* &= \left(-\frac{i}{2}f_2^* + \frac{1}{12}g_2^*\right) K_m^*\eta_1 \\
 &+ \left\{ \left(f_1 - f + \frac{i}{8}g_1 + if_0\right) \tilde{J}_{mn} + \frac{1}{2} \left( if + if_1 - \frac{i}{2}f_3 - \frac{1}{8}g_1 + \frac{1}{12}g_3 + f_0 \right) K_m^*K_n \right\} \gamma^n\eta_1^* \\
 e^\phi F\gamma_m\eta_2^* &= \left( if + if_1 + \frac{i}{2}f_3 - \frac{1}{8}g_1 - \frac{1}{12}g_3 + f_0 \right) K_m^*\eta_1 \\
 &+ \frac{1}{2} \left\{ \left(f_1 - f + \frac{i}{8}g_1 + if_0\right) \omega_{mn} + \left(-\frac{i}{2}f_2 + \frac{1}{12}g_2\right) K_m^*K_n \right\} \gamma^n\eta_1^*,
 \end{aligned} \tag{A.4}$$

for type IIA, while:

$$\begin{aligned}
e^\phi F\eta_1 &= -\frac{i}{3}f_2^*\eta_1^* + \left(g_1^* - ig_2^* + \frac{i}{6}f_3^*\right) K_m^* \gamma^m \eta_1 \\
e^\phi F\eta_2 &= -2 \left(g_1^* - ig_2^* - \frac{i}{6}f_3^*\right) \eta_1^* + \frac{i}{6}f_1^* K_m^* \gamma^m \eta_1 \\
e^\phi F\gamma_m \eta_1 &= 2 \left(g_1 + ig_2 + \frac{i}{6}f_3\right) K_m \eta_1 + \left\{ (g_2^* - ig_1^*)\omega_{mn} - \frac{i}{6}f_1 K_m K_n \right\} \gamma^n \eta_1^* \\
e^\phi F\gamma_m \eta_2 &= \frac{i}{3}f_2 K_m \eta_1 - \left\{ 2(g_2^* - ig_1^*)\tilde{J}_{mn} + \left(g_1 + ig_2 - \frac{i}{6}f_3\right) K_m K_n \right\} \gamma^n \eta_1^*,
\end{aligned} \tag{A.5}$$

$$\tag{A.6}$$

for type IIB.

The relations above can be put in a slightly different form, which is sometimes more convenient, by making use of the identities:

$$\begin{aligned}
\gamma_m \eta_1 &= -\frac{i}{2}\omega_{mn}\gamma^n \eta_2 + K_m \eta_2^* \\
&= -i\tilde{J}_{mn}\gamma^n \eta_1 + K_m \eta_2^*
\end{aligned} \tag{A.7}$$

and

$$\begin{aligned}
\gamma_m \eta_2 &= -\frac{i}{2}\omega_{mn}^*\gamma^n \eta_1 - K_m \eta_1^* \\
&= i\tilde{J}_{mn}\gamma^n \eta_2 - K_m \eta_1^*,
\end{aligned} \tag{A.8}$$

which follow from the formulæ of [16]. Taking the above into account we rewrite (A.1), (A.2) equivalently as:

$$\begin{aligned}
H\eta_1 &= \frac{i}{3}(-h_2^*\eta_1^* + h_3^*\eta_2^*) \\
H\eta_2 &= \frac{i}{3}(h_3^*\eta_1^* + h_1^*\eta_2^*)
\end{aligned} \tag{A.9}$$

and

$$\begin{aligned}
H_m \eta_1 &= \frac{i}{6}(h_3 K_m \eta_1 + h_1 K_m \eta_2 - h_2^* \gamma_m \eta_1^* + h_3^* \gamma_m \eta_2^*) \\
H_m \eta_2 &= \frac{i}{6}(h_2 K_m \eta_1 - h_3 K_m \eta_2 + h_3^* \gamma_m \eta_1^* + h_1^* \gamma_m \eta_2^*),
\end{aligned} \tag{A.10}$$

and similarly for (A.3)–(A.6).

## B Tensor decompositions

For the tensor decompositions of the various fields with respect to the local SU(2) structure we follow closely [16], to which the reader is referred for further details. In the case of the scalar ansatz the various formulæ simplify considerably, and are listed in eqs. (2.12)–(2.14).

In terms of the local SU(2) structure, the form fields decompose in general as follows.

*Two-form*

$$e^\phi F_{mn} = f_{mn} + f_{[m} K_{n]} + f_{[m}^* K_{n]}^* + i f_1 K_{[m} K_{n]}^*, \tag{B.1}$$

with

$$K^i f_{im} = K^i f_i = K^{*i} f_i = 0, \quad (\text{B.2})$$

where  $f_1$  is real. We further decompose

$$f_{mn} = \tilde{f}_{mn} + \frac{1}{8}\omega_{mn}f_2 + \frac{1}{8}\omega_{mn}^*f_2^* + \frac{1}{4}\tilde{J}_{mn}f_3, \quad (\text{B.3})$$

where  $\tilde{f}_{mn}$  is (1, 1) and traceless with respect to  $\tilde{J}_{mn}$ , i.e. it transforms in the **3** of SU(2). The scalar  $f_2$  is complex whereas  $f_3$  is real. Moreover,

$$f_m = -\frac{1}{4}\omega_m^i \tilde{f}_{1i} - \frac{1}{4}\omega_m^*{}^i \tilde{f}_{2i}, \quad (\text{B.4})$$

where  $(\Pi^-)_m{}^n \tilde{f}_{1n} = (\Pi^+)_m{}^n \tilde{f}_{2n} = 0$ . i.e.  $\tilde{f}_{1i}$  transforms in the **2** of SU(2) whereas  $\tilde{f}_{2i}$  transforms in the  $\bar{\mathbf{2}}$ .

*Three-form*

$$H_{mnp} = h_{mnp} + h_{[mn}K_{p]} + h_{[mn}^*K_{p]}^* + ih_{[m}K_nK_{p]}^*, \quad (\text{B.5})$$

with

$$K^i h_{imn} = K^i h_{im} = K^{*i} h_{im} = K^i h_i = 0, \quad (\text{B.6})$$

where  $h_m$  is real and  $h_{mn}$  is complex. We further decompose

$$h_{mnp} = -\frac{3}{32}\omega_{[mn}^*\omega_{p]}^i \tilde{h}_{1i} - \frac{3}{32}\omega_{[mn}\omega_{p]}^*{}^i \tilde{h}_{1i}^*, \quad (\text{B.7})$$

where  $(\Pi^-)_m{}^n \tilde{h}_{1n} = 0$ . Moreover

$$h_{mn} = \tilde{h}_{mn} + \frac{1}{8}\omega_{mn}h_1 + \frac{1}{8}\omega_{mn}^*h_2 + \frac{1}{4}\tilde{J}_{mn}h_3, \quad (\text{B.8})$$

where  $\tilde{h}_{mn}$  is complex and (1, 1) and traceless with respect to  $\omega_{mn}$ . The scalars  $h_{1,2,3}$  are complex. Finally,

$$h_m = -\frac{1}{4}\omega_m^i \tilde{h}_{2i} - \frac{1}{4}\omega_m^*{}^i \tilde{h}_{2i}^*, \quad (\text{B.9})$$

where  $(\Pi^-)_m{}^n \tilde{h}_{2n} = 0$ .

*Four-form*

$$e^\phi F_{mnpq} = g_{mnpq} + g_{[mnp}K_{q]} + g_{[mnp}^*K_{q]}^* + ig_{[mn}K_pK_{q]}^*, \quad (\text{B.10})$$

with

$$K^i g_{imnp} = K^i g_{imn} = K^{*i} g_{imn} = K^i g_{im} = 0, \quad (\text{B.11})$$

where  $g_{mnpq}$ ,  $g_{mn}$  are real and  $g_{mnp}$  is complex. We further decompose

$$g_{mnpq} = \frac{3}{8}\omega_{[mn}\omega_{pq]}g_1, \quad (\text{B.12})$$

where the scalar  $g_1$  is real. Moreover

$$g_{mnp} = -\frac{3}{32}\omega_{[mn}^*\omega_p]^i\tilde{g}_{1i} - \frac{3}{32}\omega_{[mn}\omega_p^*]^i\tilde{g}_{2i}, \quad (\text{B.13})$$

where  $(\Pi^-)_m{}^n\tilde{g}_{1n} = (\Pi^+)_m{}^n\tilde{g}_{2n} = 0$ . Finally,

$$g_{mn} = \tilde{g}_{mn} + \frac{1}{8}\omega_{mn}g_2 + \frac{1}{8}\omega_{mn}^*g_2^* + \frac{1}{4}\tilde{J}_{mn}g_3, \quad (\text{B.14})$$

where  $\tilde{g}_{mn}$  is real and it is traceless with respect to  $\omega_{mn}$ . The scalar  $g_2$  is complex whereas  $g_3$  is real.

*Six-form*

$$e^\phi F_{mnpqrs} = f\varepsilon_{mnpqrs}. \quad (\text{B.15})$$

For the tensor decompositions in IIB one proceeds in an analogous fashion.

### C Local SU(2) structure

This appendix contains details of the derivation of eqs. (2.26). Moreover, at the end of the section we give the torsion classes of the SU(3) structure specified by  $(J^{(1)}, \Omega^{(1)})$ . A similar computation could be used to derive the torsion classes of the SU(3) structure specified by  $(J^{(2)}, \Omega^{(2)})$ .

Plugging the tensor decompositions (2.12)–(2.14) into the differential equations (2.6), taking the formulæ in appendix A into account, we obtain:

$$\begin{aligned} \nabla_m \eta_1 &= -\partial_m \log a \eta_1 \\ &+ \{K_m(K)_1 + K_m^*(K^*)_1\} \eta_1 \\ &+ \left\{ \tilde{J}_{mn}(\tilde{J})_1 + \omega_{mn}(\omega)_1 + K_m^*K_n(K^*K)_1 + K_mK_n(KK)_1 \right\} \gamma^n \eta_1^* \end{aligned} \quad (\text{C.1})$$

and

$$\begin{aligned} \nabla_m \eta_2 &= \frac{c}{b} \partial_m \log \frac{a}{c} \eta_1 - \partial_m \log b \eta_2 \\ &+ \{K_m(K)_2 + K_m^*(K^*)_2\} \eta_1 \\ &+ \left\{ \tilde{J}_{mn}(\tilde{J})_2 + \omega_{mn}(\omega)_2 + K_m^*K_n(K^*K)_2 + K_mK_n(KK)_2 \right\} \gamma^n \eta_1^*, \end{aligned} \quad (\text{C.2})$$

where

$$\begin{aligned} (K)_1 &:= -\frac{i}{24}h_3 \\ (K^*)_1 &:= -\frac{i}{24}h_3^* + \frac{b}{8a} \left( if + if_1 + \frac{i}{2}f_3 - \frac{1}{8}g_1 - \frac{1}{12}g_3 + f_0 \right) + \frac{c^*}{8a} \left( -\frac{i}{2}f_2^* + \frac{1}{12}g_2^* \right) \\ (\tilde{J})_1 &:= -\frac{1}{24}h_2^* + \frac{c^*}{8a} \left( f_1 - f + if_0 + \frac{i}{8}g_1 \right) \\ (\omega)_1 &:= \frac{1}{48}h_3^* + \frac{b}{16a} \left( f_1 - f + if_0 + \frac{i}{8}g_1 \right) \\ (K^*K)_1 &:= +\frac{i}{48}h_2^* + \frac{b}{8a} \left( -\frac{i}{4}f_2 + \frac{1}{24}g_2 \right) + \frac{c^*}{16a} \left( if + if_1 - \frac{i}{2}f_3 - \frac{1}{8}g_1 + \frac{1}{12}g_3 + f_0 \right) \\ (KK)_1 &:= \frac{i}{48}h_1 \end{aligned} \quad (\text{C.3})$$



and

$$\begin{aligned}
(K)_2 &:= \frac{i}{24}h_2 + \frac{ic}{12b}h_3 \\
(K^*)_2 &:= \frac{i}{24}h_1^* + \frac{ic}{12b}h_3^* - \frac{c}{8a} \left( if + if_1 + \frac{i}{2}f_3 - \frac{1}{8}g_1 - \frac{1}{12}g_3 + f_0 \right) \\
&\quad + \frac{|c|^2}{8ab} \left( \frac{i}{2}f_2^* - \frac{1}{12}g_2^* \right) - \frac{a}{8b} \left( \frac{i}{2}f_2^* + \frac{1}{12}g_2^* \right) \\
(\tilde{J})_2 &:= \frac{c}{12b}h_2^* - \frac{1}{24}h_3^* - \frac{|c|^2}{8ab} \left( f_1 - f + if_0 + \frac{i}{8}g_1 \right) - \frac{a}{8b} \left( f - f_1 + if_0 + \frac{i}{8}g_1 \right) \\
(\omega)_2 &:= -\frac{1}{48}h_1^* - \frac{c}{24b}h_3^* - \frac{c}{16a} \left( f_1 - f + if_0 + \frac{i}{8}g_1 \right) \\
(K^*K)_2 &:= \frac{i}{48}h_3^* - \frac{ic}{24b}h_2^* - \frac{c}{8a} \left( -\frac{i}{4}f_2 + \frac{1}{24}g_2 \right) - \frac{|c|^2}{16ab} \left( if + if_1 - \frac{i}{2}f_3 - \frac{1}{8}g_1 + \frac{1}{12}g_3 + f_0 \right) \\
&\quad + \frac{a}{16b} \left( if + if_1 - \frac{i}{2}f_3 + \frac{1}{8}g_1 - \frac{1}{12}g_3 - f_0 \right) \\
(KK)_2 &:= -\frac{ic}{24b}h_1 + \frac{i}{48}h_3
\end{aligned} \tag{C.4}$$

for type IIA. Similarly for IIB we have:

$$\begin{aligned}
(K)_1 &:= -\frac{i}{24}h_3 - \frac{ib}{24a}f_2 - \frac{c}{4a} \left( g_1 + ig_2 + \frac{i}{6}f_3 \right) \\
(K^*)_1 &:= -\frac{i}{24}h_3^* \\
(\tilde{J})_1 &:= -\frac{1}{24}h_2^* + \frac{b}{4a}(g_2^* - ig_1^*) \\
(\omega)_1 &:= \frac{1}{48}h_3^* - \frac{c}{8a}(g_2^* - ig_1^*) \\
(K^*K)_1 &:= \frac{i}{48}h_2^* \\
(KK)_1 &:= \frac{i}{48}h_1 + \frac{b}{8a} \left( g_1 + ig_2 - \frac{i}{6}f_3 \right) + \frac{ic}{48a}f_1
\end{aligned} \tag{C.5}$$

and

$$\begin{aligned}
(K)_2 &:= \frac{i}{24}h_2 + \frac{ic}{12b}h_3 + \frac{ic}{24a}f_2 + \frac{c^2 + a^2}{4ab}(g_1 + ig_2) + \frac{i(c^2 - a^2)}{24ab}f_3 \\
(K^*)_2 &:= \frac{i}{24}h_1^* + \frac{ic}{12b}h_3^* \\
(\tilde{J})_2 &:= \frac{c}{12b}h_2^* - \frac{1}{24}h_3^* - \frac{c}{4a}(g_2^* - ig_1^*) \\
(\omega)_2 &:= -\frac{1}{48}h_1^* - \frac{c}{24b}h_3^* + \frac{c^2 + a^2}{8ab}(g_2^* - ig_1^*) \\
(K^*K)_2 &:= \frac{i}{48}h_3^* - \frac{ic}{24b}h_2^* \\
(KK)_2 &:= -\frac{ic}{24b}h_1 + \frac{i}{48}h_3 - \frac{c}{8a} \left( g_1 + ig_2 - \frac{i}{6}f_3 \right) - \frac{i(c^2 - a^2)}{48ab}f_1 .
\end{aligned} \tag{C.6}$$

It is now straightforward to read off the action of the exterior differential on the local structure. Plugging eqs. (C.1), (C.2) into the definitions (2.8), (2.9), (2.11), taking (2.10) into account, we find:

$$dK = K^* \wedge K \left\{ (K^*)_1 - 2(K^*K)_2 - \frac{1}{2}K \cdot \partial \log(ab) \right\} + \omega \{-4(\omega)_2\} + \omega^* \{2(\tilde{J})_1\} + \tilde{J} \{-2(\tilde{J})_2 - 4(\omega)_1\} . \quad (\text{C.7})$$

$$d\tilde{J} = K \wedge \omega \left\{ -2(K^*K)_1^* - i(\tilde{J})_1^* \right\} + K \wedge \omega^* \{-2(KK)_1 - 2i(\omega)_2^*\} + K \wedge \tilde{J} \left\{ (K)_1 + (K^*)_1^* + 2i(\omega)_1^* - i(\tilde{J})_2^* - K^* \cdot \partial \log a \right\} + \text{c.c.} . \quad (\text{C.8})$$

$$d\omega = K \wedge \tilde{J} \left\{ 2(KK)_1 + (K^*)_2^* + 4i(\omega)_2^* - \frac{c^*}{2b}K^* \cdot \partial \log \frac{c^*}{a} \right\} + K^* \wedge \tilde{J} \left\{ 2(K^*K)_1 + (K)_2^* - 2i(\tilde{J})_1 - \frac{c^*}{2b}K \cdot \partial \log \frac{c^*}{a} \right\} + K \wedge \omega \left\{ (K)_1 - 2(K^*K)_2^* - 2i(\tilde{J})_2^* - \frac{1}{2}K^* \cdot \partial \log(ab) \right\} + K^* \wedge \omega \left\{ (K^*)_1 - 2(KK)_2^* - 4i(\omega)_1 - \frac{1}{2}K \cdot \partial \log(ab) \right\} . \quad (\text{C.9})$$

The content of the three equations above is exactly equivalent to the content of the spinorial equations (C.1), (C.2). Moreover we have:

$$dJ^{(1)} = K \wedge \omega \left\{ -2(K^*K)_1^* - 2i(\tilde{J})_1^* \right\} + K \wedge \omega^* \{-2(KK)_1\} + K \wedge \tilde{J} \left\{ (K)_1 + (K^*)_1^* + 4i(\omega)_1^* - K^* \cdot \partial \log a \right\} + \text{c.c.} . \quad (\text{C.10})$$

$$d\Omega^{(1)} = K^* \wedge K \wedge \tilde{J} \left\{ -4i(K^*K)_1 - 2(\tilde{J})_1 \right\} + K^* \wedge \Omega^{(1)} \left\{ 2(K^*)_1 - 4i(\omega)_1 - K \cdot \partial \log a \right\} + \tilde{J} \wedge \tilde{J} \left\{ -4i(\tilde{J})_1 \right\} . \quad (\text{C.11})$$

$$dJ^{(2)} = K \wedge \omega \left\{ (K)_2 - \frac{c}{2b}K^* \cdot \partial \log \frac{c}{a} \right\} + K \wedge \omega^* \left\{ (K^*)_2^* + 4i(\omega)_2^* - \frac{c^*}{2b}K^* \cdot \partial \log \frac{c^*}{a} \right\} + K \wedge \tilde{J} \left\{ 2(KK)_2 + 2(K^*K)_2^* + 2i(\tilde{J})_2^* + K^* \cdot \partial \log b \right\} + \text{c.c.} . \quad (\text{C.12})$$

$$d\Omega^{(2)} = K^* \wedge K \wedge \tilde{J} \left\{ 2i(K^*)_2 + 4(\omega)_2 - \frac{ic}{b}K \cdot \partial \log \frac{c}{a} \right\} + K^* \wedge K \wedge \omega^* \left\{ -4i(K^*K)_2 - 2(\tilde{J})_2 - iK \cdot \partial \log b \right\} + \tilde{J} \wedge \tilde{J} \{-8i(\omega)_2\} . \quad (\text{C.13})$$

It is also useful to define:

$$\tilde{\Omega}_{mnp} := \eta_2 \gamma_{mnp} \eta_1 , \quad (\text{C.14})$$

so that:

$$\tilde{\Omega} = i\tilde{J} \wedge K . \quad (\text{C.15})$$

We find:

$$\begin{aligned}
d\tilde{\Omega} = & K^* \wedge K \wedge \tilde{J} \left\{ i(K^*)_1 - 2i(K^*K)_2 + 2(\omega)_1 - (\tilde{J})_2 - \frac{i}{2}K \cdot \partial \log(ab) \right\} \\
& + K^* \wedge \Omega^{(1)} \left\{ (K^*)_2 - 2i(\omega)_2 - \frac{c}{2b}K \cdot \partial \log \frac{c}{a} \right\} \\
& + \tilde{J} \wedge \tilde{J} \left\{ -4i(\omega)_1 - 2i(\tilde{J})_2 \right\} + K^* \wedge K \wedge \omega^* \left\{ -2i(K^*K)_1 - (\tilde{J})_1 \right\} . \quad (C.16)
\end{aligned}$$

One can perform several consistency checks of these expressions. For example,  $d\Omega^{(1)}$  can be computed in two different ways: either directly by plugging eq. (C.1) into definition (2.8), or by plugging the expressions for  $d\omega$ ,  $dK$  above into  $d\Omega^{(1)} = -id\omega \wedge K - i\omega \wedge dK$ , which follows from eq. (2.10). In order to perform these consistency checks, it is useful to take the following equations into account:

$$\begin{aligned}
(K)_1 + (K^*)_1^* &= K^* \cdot \partial \log a \\
(K^*K)_2 + (KK)_2^* &= -\frac{1}{2}K \cdot \partial \log b \\
(K^*)_2 - 2(KK)_1^* &= \frac{c}{2b}K \cdot \partial \log \frac{c}{a} \\
(K)_2 - 2(K^*K)_1^* &= \frac{c}{2b}K^* \cdot \partial \log \frac{c}{a} \\
d \log \frac{|c|}{a} &= -\frac{b^2}{|c|^2} d \log \frac{b}{a} . \quad (C.17)
\end{aligned}$$

The first four equations above can be shown to be equivalent to  $\mathcal{L}_K(\eta_i^\dagger \eta_j) = \mathcal{L}_{K^*}(\eta_i^\dagger \eta_j) = 0$ , for  $i, j = 1, 2$ , once (C.1), (C.2) are taken into account. The last relation follows from (2.22). Alternatively, eqs. (C.17) can be derived directly from the solution (2.15), (2.22) and the constraints (2.16) in IIA, and similarly in IIB.

**Torsion classes.** As discussed in some detail in section 2, each of the two spinors  $\theta_{1,2}$  can be used to define an SU(3) structure on  $\mathcal{M}_6$ . On the other hand, for an SU(3)-structure manifold, the torsion classes are defined via:

$$\begin{aligned}
dJ &= \frac{3i}{4} (\mathcal{W}_1 \Omega^* - \mathcal{W}_1^* \Omega) + \mathcal{W}_3 + \mathcal{W}_4 \wedge J \\
d\Omega &= \mathcal{W}_1 J \wedge J + \mathcal{W}_2 \wedge J + \mathcal{W}_5^* \wedge \Omega . \quad (C.18)
\end{aligned}$$

In particular, the torsion classes corresponding to the SU(3) structure  $(J^{(1)}, \Omega^{(1)})$  can be read off by comparing the above with (C.10), (C.11), taking (C.17) into account:

$$\begin{aligned}
\mathcal{W}_1 &= -\frac{8i}{3} \left\{ (\tilde{J})_1 + i(K^*K)_1 \right\} \\
\mathcal{W}_2 &= -\frac{4i}{3} \left( J^{(1)} - \frac{3i}{2}K \wedge K^* \right) \left\{ (\tilde{J})_1 - 2i(K^*K)_1 \right\} \\
\mathcal{W}_3 &= K \wedge \omega^* \left\{ -2(KK)_1 \right\} + \text{c.c.} \\
\mathcal{W}_4 &= K \left\{ 4i(\omega)_1^* \right\} + \text{c.c.} \\
\mathcal{W}_5 &= K \left\{ (K^*)_1^* - (K)_1 + 4i(\omega)_1^* \right\} . \quad (C.19)
\end{aligned}$$

$\mathcal{M}$	$\mathcal{Hol}(\mathcal{C}(\mathcal{M}_7))$	$\mathcal{N}$
Weak $G_2$	Spin(7)	1
Sasaki-Einstein	SU(4)	2
tri-Sasaki	Sp(2)	3
$S^7$	1	8

**Table 1.** List of seven-dimensional Einstein manifolds  $\mathcal{M}_7$ , the holonomy of the corresponding eight-dimensional cones and the number of preserved supersymmetries in four dimensions.

We see that  $\mathcal{W}_5$  is proportional to  $K$ . Moreover, in the IIA case, taking (2.15), (C.3) into account we find that  $\mathcal{W}_4$  is exact:  $\mathcal{W}_4 = d(\phi - 3A)$ . Therefore  $\mathcal{W}_4$  can be removed by a conformal rescaling of the internal metric:  $ds_6^2 \rightarrow e^{3A-\phi} ds_6^2$ .

### D Sasaki-Einstein

There is a well-known class of eleven-dimensional supergravity solutions of the form  $AdS_4 \times \mathcal{M}_7$ , where  $\mathcal{M}_7$  is a seven-dimensional Einstein manifold. Specifically, the eleven - dimensional metric is given by

$$ds^2 = ds^2(AdS_4) + ds^2(\mathcal{M}_7), \tag{D.1}$$

while the four-form flux is of Freund-Rubin type:  $G_4 \propto vol_4$ , where  $vol_4$  is the volume form of  $AdS_4$ . In addition, the manifold  $\mathcal{M}_7$  has the property that the cone over it,  $\mathcal{C}(\mathcal{M}_7)$ , is an eight-dimensional manifold of special holonomy. The supersymmetry preserved by the solution depends on the holonomy of  $\mathcal{C}(\mathcal{M}_7)$ . Table 1 lists the type of the seven-dimensional Einstein manifold  $\mathcal{M}_7$ , the holonomy of the cone over it,  $\mathcal{Hol}(\mathcal{C}(\mathcal{M}_7))$ , as well as the number of preserved supersymmetries,  $\mathcal{N}$ , in four dimensions.

We will now specialize to the case where  $\mathcal{Hol}(\mathcal{C}(\mathcal{M}_7))$  is a subgroup of SU(4), i.e. the eight-dimensional cone is Calabi-Yau. Equivalently, we will take  $\mathcal{M}_7$  to be Sasaki-Einstein (which includes the  $S^7$  and the tri-Sasaki as special cases). The manifold  $\mathcal{M}_7$  can then be thought of as the total space of a fibre bundle with connection one-form  $\mathcal{A}$  on a six-dimensional base-space  $\mathcal{M}_6$ ,

$$ds^2(\mathcal{M}_7) = (dy + \mathcal{A})^2 + ds^2(\mathcal{M}_6), \tag{D.2}$$

where  $ds^2(\mathcal{M}_6)$  is a local Kähler-Einstein metric and  $y$  is the coordinate on the fibre. The Killing vector  $\partial_y$  is the so-called ‘Reeb vector’. If the orbits of the Reeb vector are closed and the U(1) action is free,  $\mathcal{M}_7$  is regular and  $\mathcal{M}_6$  is globally a manifold. One can define a local SU(3) structure on  $\mathcal{M}_6$  specified by a Kähler form  $J$  and a complex three-form  $\Omega$ , such that  $d\mathcal{A} = 2J$  and  $d\Omega = 4i\mathcal{A} \wedge \Omega$ . Note, however, that globally the structure group of  $\mathcal{M}_6$  is not SU(3) but rather U(3), since  $\Omega$  is not globally defined in general.

A useful property of odd-dimensional, simply-connected Sasaki-Einstein manifolds is that they admit at least two Killing spinors. In the seven-dimensional case, it was shown in [28] that, under certain regularity assumptions, the converse is also true: any pair of

(real) Killing spinors defines a Sasaki-Einstein structure on  $\mathcal{M}_7$ . Moreover, there is a one-to-one correspondence between triplets of Killing spinors and tri-Sasaki structures on  $\mathcal{M}_7$ .

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